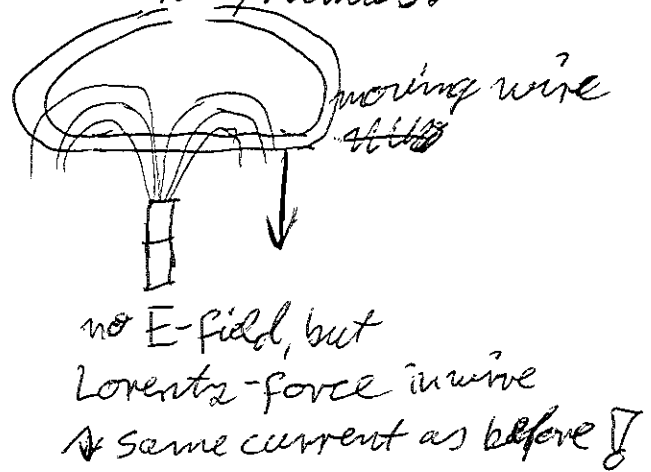
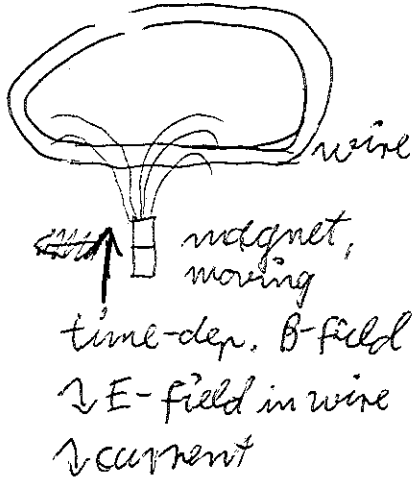


Special Relativity

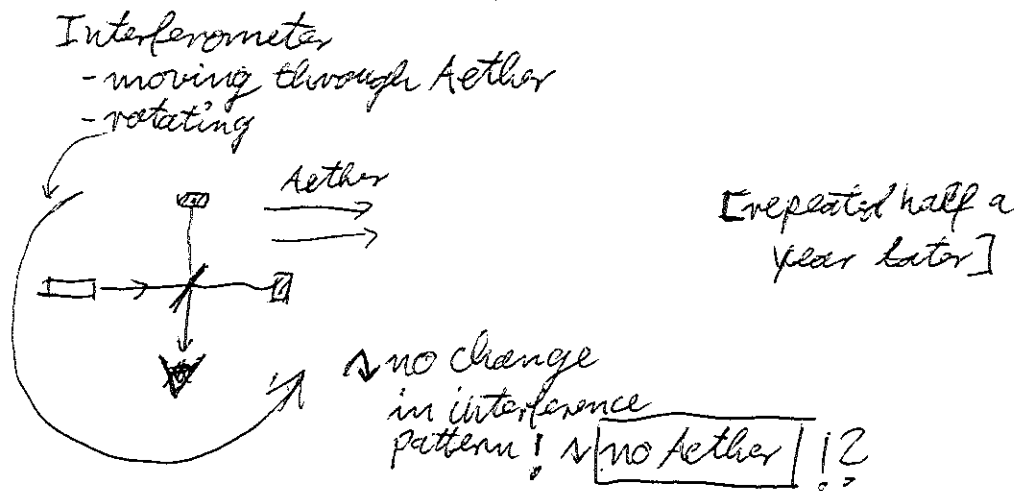
IMPRS 2018 LTP1

[Goal: spacetime geometry! History...
new theory can't be deduced & motivation possible]

Einstein (1905): "On the electrodynamics of moving bodies"
unnatural asymmetry in electrodynamics:



Also: Michelson-Morley experiment



↳ Principle of special relativity!
in all inertial systems, where the mechanical laws hold, the electrodynamic equations also hold

↳ no Aether

↳ speed of light c in all inertial systems is the same

[in GR: all laws of physics!]

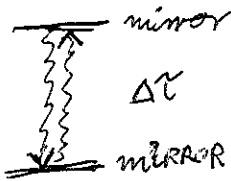
[extra assumption, could change Edyn]

useful units: $c=1$

$\Delta L \approx 3 \cdot 10^5 \text{ km} \approx 80\%$ distance Earth-Moon

time dilation time is what a clock measures

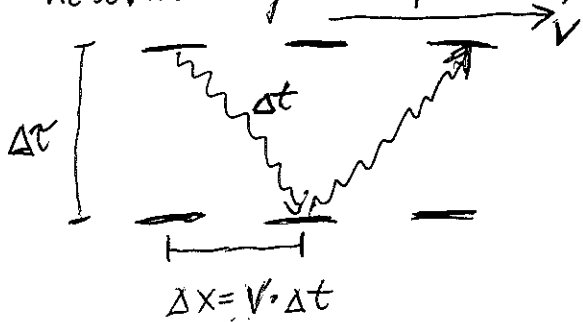
a clock based on light-travel:



1 tick of the clock = $2\Delta\tau$
 $\hat{=}$ photon travel time

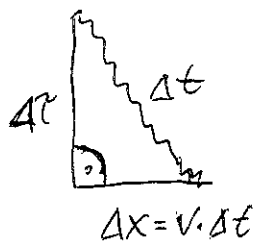
$\hat{=}$ 2 · distance between mirrors ($c=1$)

now: moving clock, velocity V



$2 \cdot \Delta\tau$: photon travel time

$\Delta t \neq \Delta\tau$ since photon travels a longer distance!



$$\Delta t^2 = \Delta\tau^2 + \underbrace{\Delta x^2}_{V^2 \Delta\tau^2}$$

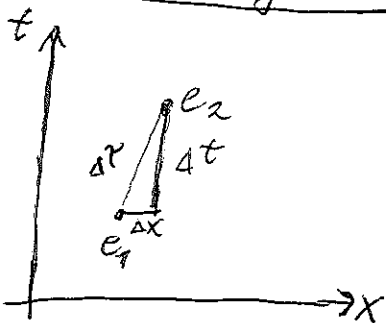
→ Important: $\Delta\tau^2 = \Delta t^2 - \Delta x^2$

$$\hookrightarrow \frac{\Delta t}{\Delta\tau} = \gamma = \frac{1}{\sqrt{1-V^2}} \geq 1 \text{ "gamma factor"}$$

$\hookrightarrow \Delta t \geq \Delta\tau$ time dilation

spacetime diagrams

[careful: before XY-plane, now xt-plane]



events: points in spacetime

e_1 : breakfast on Earth

e_2 : dinner on the moon

$\Delta\tau$: travel time in rest-system

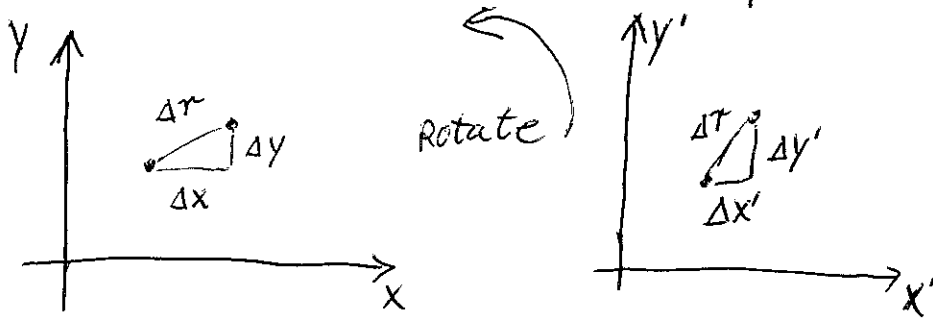
= proper time (between e_1 and e_2)

$\Delta\tau^2 = \Delta t^2 - \Delta x^2 = \text{invariant}$

- $\Delta x, \Delta t$ are observer-dependent
- another observer measures $\Delta x', \Delta t'$

but: $\Delta \tau^2 \equiv \Delta t^2 - \Delta x^2 = \Delta t'^2 - \Delta x'^2 = \text{invariant} \forall$

analogous to distance in Euclidean space:



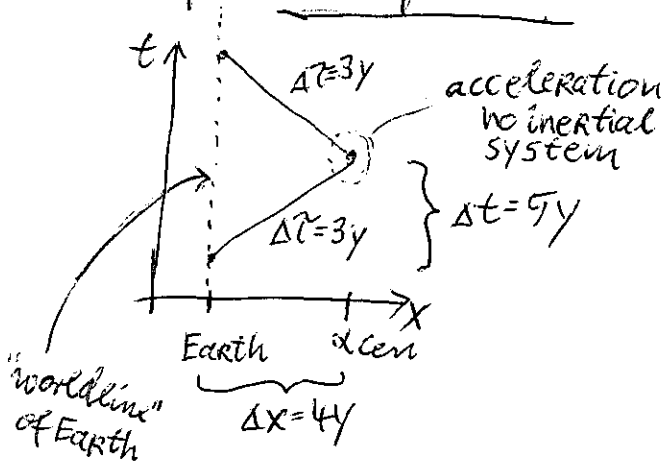
$$\begin{aligned} \Delta r^2 &= \Delta x'^2 + \Delta y'^2 \\ &= \Delta x'^2 + \Delta y'^2 \\ &= \text{invariant under} \\ &\text{Rotation of} \\ &\text{coordinate system} \end{aligned}$$

$\Delta \tau$ is an invariant "distance" between events in spacetime \forall

in general: $\Delta \tau^2 = \Delta t^2 - \Delta x^2 - \Delta y^2 - \Delta z^2$ "spacetime-Pythagoras" line element

↑ ↑ ↑
minus \forall

example: twin paradox



$$\begin{aligned} \Delta \tau^2 &= \Delta t^2 - \Delta x^2 \\ &= 25 - 16 = 9 \end{aligned}$$

$\Delta \tau = 3y$ \rightarrow time in spaceship running slower!

$$\Delta \tau^2 = \Delta t'^2 - \Delta x'^2$$

$\Delta \tau = \Delta t' = \text{proper time}$

paradox: seen from the spaceship, time on Earth is running slower.

Resolution: during acceleration, time on Earth must catch up.

the geometric point of view gives clear answers \rightarrow no paradox \forall

advice: draw spacetime diagram in inertial frames

Remark: acceleration reduces proper time

Lorentz-contraction

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How can the spaceship travel $\Delta x = 4y$ in $\Delta t = 3y$?

↳ distance Earth- α -cen must be contracted as seen from the spaceship

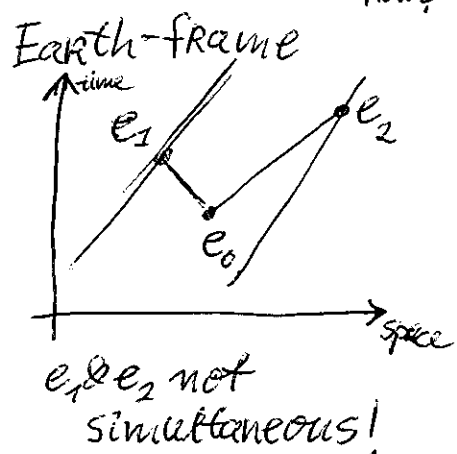
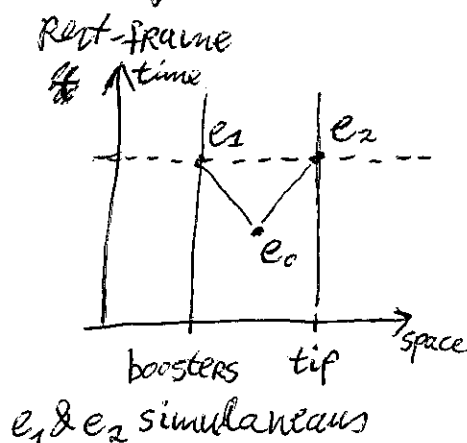
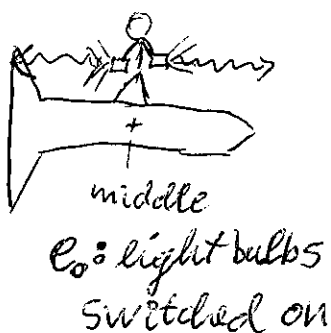
$$\Delta t' = \frac{\Delta t}{\gamma}, \text{ distance seen from spaceship} = L', \Delta x = L$$

$$L' = v \cdot \Delta t' = \frac{v \cdot \Delta t}{\gamma} = \frac{L}{\gamma}$$

$\gamma \geq 1 \Rightarrow L' \leq L$ length contraction

Relativity of simultaneity

[Imagine L' is contracted to length of the ship. Ship contracted seen from Earth? How?]



proper length Δs : length measured in rest-frame of the object

abstract: proper length between two events

= length measured in system where events are simultaneous

we argued: $\Delta t^2 - \Delta x^2 = \text{invariant}$

in rest-frame of an object: $\Delta t = 0$ (simultaneous)

and $\Delta x = \Delta s$

$$\Rightarrow \Delta s^2 = \Delta x^2 - \Delta t^2 = \text{invariant}$$

conclusions:

events: points in spacetime

invariant "distance" between events:

line element
$$\Delta t^2 - \Delta x^2 = \begin{cases} \Delta \tau^2 & \text{for } |\Delta t| > |\Delta x| \\ -\Delta s^2 & \text{for } |\Delta t| < |\Delta x| \end{cases}$$

↳ spacetime geometry ⚡