

Intro:  $\mu, \nu, \dots = 0, 1, 2, 3$  } indices  
 $i, j, \dots = 1, 2, 3$

$$\eta = \text{diag}(-1, 1, 1, 1) \quad G = I = C$$

## Gravitational waves (GWs)

max. speed: c

↳ no instantaneous forces

↳ propagation, waves ↳

↳ electromag., gravitational

particle physics

↳ relativity + quantum mechanics

↳ forces ↳ bosonic particles ↳ waves (duality)

graviton: massless spin 2 particle

↳ 2 polarizations!

## Linearized gravity

$$g_{\mu\nu} = \eta_{\mu\nu} + h_{\mu\nu}, \quad |h_{\mu\nu}| \ll 1$$

↑ flat spacetime

Indices now pulled with  $\eta^{\mu\nu}$

$$\text{in } R_{\mu\nu} - \frac{1}{2} g_{\mu\nu} R = 8\pi T_{\mu\nu}$$

(exercise 3.1)

$$\sim \cancel{\partial_\alpha \partial^\alpha} \bar{h}_{\mu\nu} - \partial^\alpha \partial_\mu \bar{h}_{\nu\alpha} - \partial^\alpha \partial_\nu \bar{h}_{\mu\alpha} + \eta_{\mu\nu} \partial^\alpha \partial^\beta \bar{h}_{\alpha\beta} = -16\pi T_{\mu\nu}$$

$$\square \quad \text{where } \bar{h}_{\mu\nu} = h_{\mu\nu} - \frac{1}{2} \eta_{\mu\nu} h^\alpha_\alpha \quad \text{"trace-reversal"}$$

gauge/coordinate fixing:  $\partial^\alpha \bar{h}_{\mu\alpha} = 0$

$$\sim \boxed{\square \bar{h}_{\mu\nu} = -16\pi T_{\mu\nu}}$$

vacuum  $T_{\mu\nu}=0$   $\Rightarrow$  plane-wave solutions

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transverse-traceless (TT)-gauge for  
monochromatic plane waves;  
e.g. wave in z-direction

(exercise 4)

$$(h_{\mu\nu}^{TT}) = \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & h_+ & h_x & 0 \\ 0 & h_x & -h_+ & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} = \underbrace{\begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & a_+ & a_x & 0 \\ 0 & a_x & -a_+ & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix}}_{a_{ij}^{TT}} e^{ik_\mu x^\mu} + c.c. \quad (*)$$

$k_\mu$ : wave vector =  $(-\omega, \vec{k})$

makes polarizations  $+x$  manifest (physical DOF)

in general:  $h_{0\mu}^{TT}=0$ ,  $\partial^j h_{ij}^{TT}=0$ ,  $\text{TR}(h^{TT})=0$   $\nabla$  traceless

$\partial^j a_{ij}^{TT}=0$   $\nabla$  transverse

## Are GWs observable?

↳ might be a gauge effect

↳ should carry energy, but:

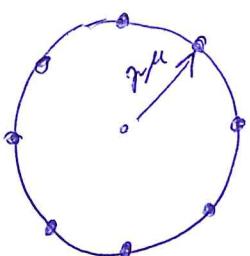
grav. energy can not be localised  
due to equivalence principle

(also: nonlinearity,  
are there exact GW sol.)

look at free-falling ring of test-masses:

geodesic deviation

$$\frac{D^2 r^\mu}{Dt^2} = R^\mu_{\alpha\beta\gamma} u^\alpha u^\beta \tau^\gamma$$



$u^\mu$ : 4-velocity of ring

$\tau$ : proper time " "

$t$ : coordinate time

after "some" calculation in linear growth:

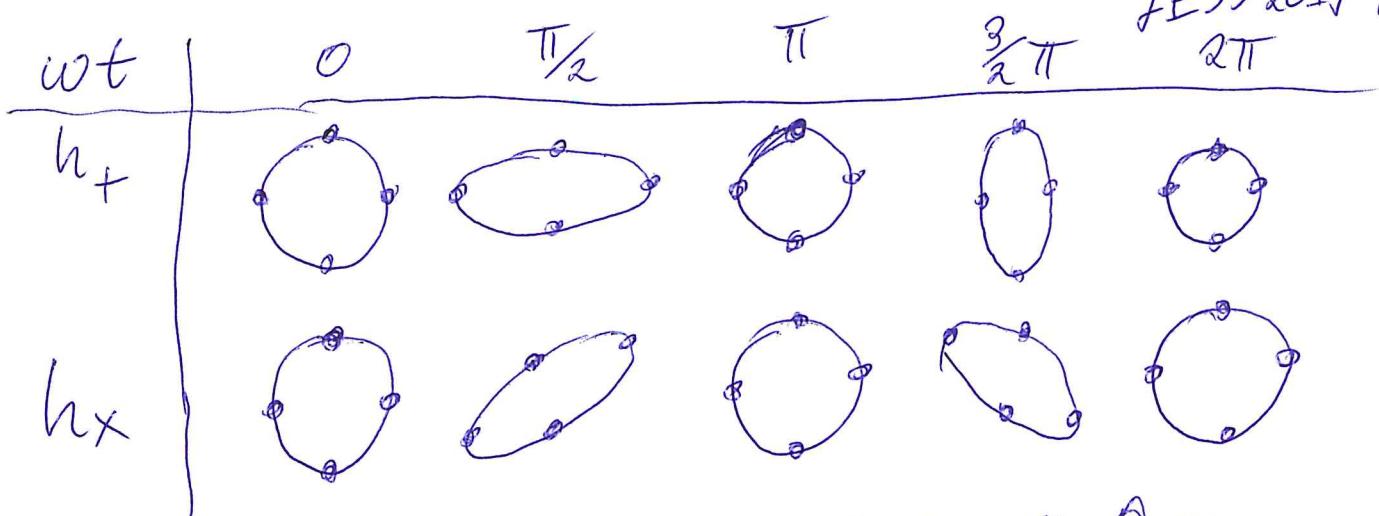
(exercise 2)

$$\frac{d^2 r_i}{dt^2} \approx \frac{1}{2} r^j \frac{\partial^2 h_{ij}^{TT}}{\partial t^2}$$

ansatz:  $r_i = r_0^i + \Delta r^i$ ,  $|\Delta \vec{r}| \ll |\vec{r}_0|$  and (\*)

$$\approx \Delta r^i \approx \frac{1}{2} r_0^i h_{ij}^{TT} = \frac{1}{2} \begin{pmatrix} 0 & 0 & 0 & 0 \\ 0 & h_+ & h_x & 0 \\ 0 & h_x & -h_+ & 0 \\ 0 & 0 & 0 & 0 \end{pmatrix} \vec{r}_0$$

&  $h_{+,x} \approx \sin \omega t$



circular pol. from superposition:  $\circ \circ \circ \circ \dots$

## GW detectable

$h \sim$  (dim. less) strain

gravity (linear)	electrodynamics
harmonic gauge $\partial^\mu \bar{h}_{\mu\nu} = 0$ Einstein eq: $\square \bar{h}_{\mu\nu} = -16\pi T_{\mu\nu}$ Solution $\bar{h}_{\mu\nu} = 4 \int d^3x' \frac{T_{\mu\nu}(x', t_{\text{ret}})}{ x - x' }$ $t_{\text{ret}} = t -  x - x' $ quadrupole radiation 2 polarizations 2 transverse-traceless waves	Lorentz gauge $\partial^\mu A_\mu = 0$ Maxwell eq: $\square A^\mu = -4\pi j^\mu$ $A^\mu = \int d^3x' \frac{j^\mu(x', t_{\text{ret}})}{ x - x' }$ dipolar rad. 2 polarization 2 transverse waves

far-zone approximation:  $|\vec{x}| \gg |\vec{x}'|$

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$$\sim |\vec{x} - \vec{x}'| \approx |\vec{x}| = R, t_{\text{ret}} \approx t - R$$

$$T_{\mu\nu} \approx \frac{4}{R} \int d^3x' T_{\mu\nu}(\vec{x}', t_{\text{ret}})$$

... some calculation, similar to edyn ...

$$\boxed{\tilde{h}_{ij}^{ij} \propto \frac{2}{R} \tilde{M}_{ij}(t_{\text{ret}})}$$

where  $\tilde{M}^{ij} = \int d^3x T^{00} X^i X^j$   
↑ energy/mass density

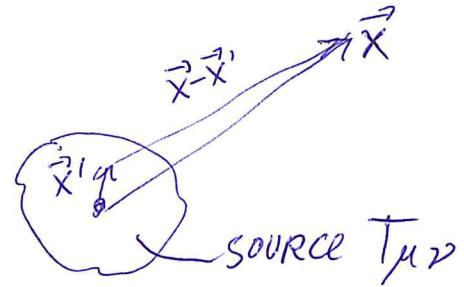
cp. quadrupole

$$Q^{ij} = \int d^3x T^{00} (X^i X^j - \frac{1}{3} S^{ij} X^k X_k)$$

$$= M^{ij} - \frac{1}{3} M^{kl} S^{ij}$$

project to TT-part: (exercise 4)

$$\boxed{\tilde{h}_{ij}^{TT} \propto \frac{2}{R} \underbrace{N_{ijkl}}_{\text{TT-projector}} \tilde{Q}^{kl}(t_{\text{ret}})}$$



quadrupole  
approximation