

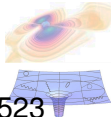
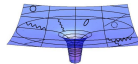
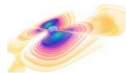
Canonical formulation of spinning objects in General Relativity

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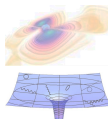
DPG Spring Meeting, March 19th, 2010, Bonn



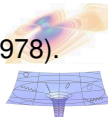
DFG: SFB/TR7 “Gravitational Wave Astronomy” and GRK 1523



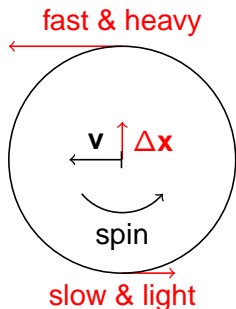
- 1 Introduction
- 2 Action Approach
- 3 Order-by-Order Construction
- 4 Results Linear in Spin
- 5 Higher Orders in Spin (Talk by S. Hergt, today 14:00)



- Why?
 - Compact objects with spin interesting for GW astronomy.
 - Canonical formalism efficient for calculating conservative dynamics.
 - Consistency check . . .
- Spinning test-bodies coupled to gravity:
 - H. P. Künzle, *J. Math. Phys.* **13**, 739 (1972).
 - K. Yee and M. Bander, *PRD* **48**, 2797 (1993).
 - E. Barausse, E. Racine, and A. Buonanno, *PRD* **80**, 104025 (2009).
- Dirac field coupled to gravity:
 - P. A. M. Dirac, in *Recent Developments in GR* (1962).
 - **T. W. B. Kibble, *J. Math. Phys.* **4**, 1433 (1963).**
 - S. Deser and C. J. Isham, *PRD* **14**, 2505 (1976).
 - J. Geheniau and M. Henneaux, *GRG* **8**, 611 (1977).
 - J. E. Nelson and C. Teitelboim, *Ann. Phys. (N.Y.)* **116**, 86 (1978).



Spin in Special Relativity



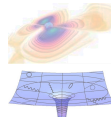
- Spin is a 4-tensor $S^{\mu\nu}$:
 - Spin is $S^{ij} = \varepsilon^{ijk} S_k$.
 - Mass dipole related to S^{i0} .
- Different mass centers.
- Need spin supplementary condition:
 - Møller SSC: $\tilde{S}^{\mu 0} = 0$
 - Covariant SSC: $S^{\mu\nu} p_\nu = 0$
 - **Newton-Wigner (canonical) SSC:**
 $m \hat{S}^{\mu 0} + \hat{S}^{\mu\nu} p_\nu = 0$

- In covariant SSC, with position \mathbf{z} :

$$\{z^i, z^j\} = \frac{S^{ij}}{m^2} - \frac{p^j S^{0j} - p^j S^{0i}}{m^2 p^0}, \quad \dots$$

- In Newton-Wigner SSC:

$$\{\hat{z}^i, p_j\} = \delta_{ij}, \quad \{\hat{S}_i, \hat{S}_j\} = \varepsilon_{ijk} \hat{S}_k$$



- Normal vector:

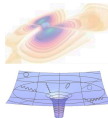
$$n_\mu = (-N, 0, 0, 0), \quad n^\mu = \frac{1}{N}(1, -N^i), \quad n_\mu n^\mu = -1$$

- Projector:

$$\gamma^{\mu\nu} = g^{\mu\nu} + n^\mu n^\nu = \begin{pmatrix} 0 & 0 \\ 0 & \gamma^{ij} \end{pmatrix}, \quad g_{ij} = \gamma_{ij}, \quad \gamma_{ik} \gamma^{kj} = \delta_{ij}$$

- Extrinsic curvature:

$$K_{ij} \equiv -n_{(i|j)}$$
$$\pi^{ij} = -\sqrt{\gamma}(\gamma^{ik} \gamma^{jl} - \gamma^{ij} \gamma^{kl}) K_{kl}$$



Point-Mass Action in ADM Form

$$W = \int dt p_i \dot{z}^i + \int d^4x \left[\frac{1}{16\pi} \pi^{ij} \gamma_{ij,0} - N \mathcal{H} + N^i \mathcal{H}_i + (\text{st}) \right]$$

- Constraint equations:

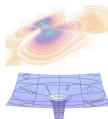
$$0 = \mathcal{H} \equiv -\frac{1}{16\pi\sqrt{\gamma}} \left[\gamma R + \frac{1}{2} \left(\gamma_{ij} \pi^{ij} \right)^2 - \gamma_{ij} \gamma_{kl} \pi^{ik} \pi^{jl} \right] + \mathcal{H}^{\text{matter}}$$

$$0 = \mathcal{H}_i \equiv \frac{1}{8\pi} \gamma_{ij} \pi^{jk}{}_{;k} + \mathcal{H}_i^{\text{matter}}$$

- **Source terms** are related to the stress-energy tensor $T^{\mu\nu}$:

$$\mathcal{H}^{\text{matter}} \equiv \sqrt{\gamma} T_{\mu\nu} n^\mu n^\nu = \sqrt{m^2 + \gamma^{ij} p_i p_j} \delta$$

$$\mathcal{H}_i^{\text{matter}} \equiv -\sqrt{\gamma} T_{i\nu} n^\nu = p_i \delta$$



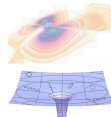
- Hamiltonian without gauge fixing:

$$H[z^i, p_i, \gamma_{ij}, \pi^{ij}] = \int d^3\mathbf{x} (N\mathcal{H} - N^i\mathcal{H}_i) + E[\gamma_{ij}]$$
$$E[\gamma_{ij}] = \frac{1}{16\pi} \oint d^2s_i (\gamma_{ij,j} - \gamma_{ji,i})$$

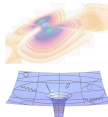
- **ADM Hamiltonian** (Hamiltonian in ADMTT gauge)
 $\hat{=}$ **ADM Energy depending on canonical variables:**

$$H_{\text{ADM}} = E[z^i, p_i, h_{ij}^{\text{TT}}, \pi_{\text{TT}}^{ij}] = -\frac{1}{16\pi} \int d^3\mathbf{x} \Delta\phi$$
$$\gamma_{ij} = \left(1 + \frac{1}{8}\phi\right)^4 \delta_{ij} + h_{ij}^{\text{TT}}, \quad \pi^{ii} = 0$$

- Matter only Hamiltonian: Elimination of h_{ij}^{TT} and π_{TT}^{ij} .



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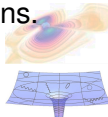


Non-Relativistic Spherical Top

- Body fixed coordinates $z^{[l]}$: $z^i(t) = R_{[l]i}(t)z^{[l]}$, $R_{[k]i}R_{[k]j} = \delta_{ij}$
- Independent angle variables: $R_{[l]j} = R_{[l]j}(\phi, \psi, \theta)$
- Angular velocity tensor: $\Omega^{ij} = \varepsilon_{ijk}\Omega^k = R_{[k]i}\dot{R}_{[k]j}$
- Lagrangian: $L = \frac{1}{4}J\Omega^{ij}\Omega^{ij}$
- Spin tensor: $S_{ij} = 2\frac{\partial L}{\partial \Omega^{ij}} = J\Omega^{ij}$
- Legendre transformed:

$$L = \frac{1}{2}S_{ij}\Omega^{ij} - H[R_{ij}, S_{ij}], \quad H = \frac{1}{4J}S_{ij}S_{ij}$$

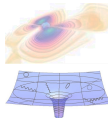
- Trick: Use $\delta\theta_{ij} = -\delta\theta_{ji} = R_{[k]i}\delta R_{[k]j}$ as independent variations.
- Usual Poisson brackets.



Relativistic Spherical Top

E.g., Hanson and Regge (1974)

- No rigid bodies.
- **Mathematical abstraction**: Top is
 - Worldline with Lorentz-matrix $\Lambda_{A\mu}$. $\eta^{AB}\Lambda_{A\mu}\Lambda_{B\nu} = \eta_{\mu\nu}$
 - $\Lambda_{A\mu}$ is pure rotation in rest-frame: $\Lambda_{A\mu} = \begin{pmatrix} -1 & 0 \\ 0 & R_{[ij]} \end{pmatrix}$
- Equivalent description: $\Lambda_{[0]\mu} = p_\mu/m$ or $\Lambda^{[i]\mu} p_\mu = 0$
- Angular velocity tensor: $\Omega^{\mu\nu} = \Lambda_A^\mu \frac{d\Lambda^{A\nu}}{d\tau}$
- Spin tensor: $S_{\mu\nu} = 2 \frac{\partial L}{\partial \Omega^{\mu\nu}}$
- Associated SSC: $S_{\mu\nu} p^\mu = 0$



Minimal Coupling

- Problem with metric variation: (also $\Lambda_{A\mu} = e_{A\mu}$)

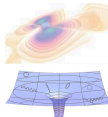
$$\Lambda_{A\mu} \Lambda^A{}_\nu = g_{\mu\nu} \quad \leftrightarrow \quad \gamma_\mu \gamma_\nu + \gamma_\nu \gamma_\mu = 2g_{\mu\nu}$$

- Vary Λ^{Aa} and tetrad $e_{a\mu}$, $\Lambda^A{}_\mu = \Lambda^{Aa} e_{a\mu}$:

$$\Lambda_{Aa} \Lambda^A{}_b = \eta_{ab} \quad \leftrightarrow \quad \gamma_a \gamma_b + \gamma_b \gamma_a = 2\eta_{ab}$$

- Matter Lagrangian density and constraints:

$$\mathcal{L}_M = \int d\tau \left[p_\mu u^\mu + \frac{1}{2} S_{ab} \Omega^{ab} \right] \delta_{(4)}$$
$$\Omega^{ab} = \Lambda_A{}^a \frac{D\Lambda^{Ab}}{d\tau} = \Lambda_A{}^a \left[\frac{d\Lambda^{Ab}}{d\tau} - \Lambda^A{}_c \omega_\mu{}^{cb} u^\mu \right]$$
$$S_{ab} p^b = 0, \quad \Lambda^{[i]a} p_a = 0, \quad p_\mu p^\mu + m^2 = 0$$



Result of Variation

- Approximated linear in spin.
- Field equations with stress-energy tensor (Mathisson, Tulczyjew):

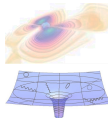
$$\sqrt{-g} T^{\mu\nu} = \int d\tau \left[m u^\mu u^\nu \delta_{(4)} - (S^{\alpha(\mu} u^{\nu)}) \delta_{(4)} \right]_{||\alpha}$$

- EOM (Mathisson, Papapetrou):

$$\frac{DS^{\mu\nu}}{d\tau} = 0, \quad \frac{Dp_\mu}{d\tau} = \frac{1}{2} S^{\lambda\nu} u^\gamma R_{\mu\gamma\nu\lambda}$$

- Original derivation: Evaluate $T^{\mu\nu}_{||\nu} = 0$ with ansatz

$$\sqrt{-g} T^{\mu\nu} = \int d\tau \left[t^{\mu\nu} \delta_{(4)} + (t^{\mu\nu\alpha} \delta_{(4)}) \right]_{||\alpha}$$



Reduction of the Matter Part

- Solve matter constraints, Schwinger time gauge $e_{(0)\mu} = n_\mu$, $\tau = t$.
- Variable redefinitions:

$$z^i = \hat{z}^i - \frac{nS^i}{m - np}, \quad np = -\sqrt{m^2 + \gamma^{ij} p_i p_j}$$

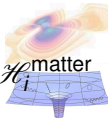
$$S_{ij} = \hat{S}_{ij} - \frac{p_i n S_j}{m - np} + \frac{p_j n S_i}{m - np}, \quad nS_i = -\frac{p_k \gamma^{kj} \hat{S}_{ji}}{m}$$

$$\Lambda^{[l](j)} = \hat{\Lambda}^{[l](k)} \left(\delta_{kj} + \frac{p_{(k} p_{j)}}{m(m - np)} \right), \quad \gamma_{ik} \gamma_{jl} A^{kl} = \frac{1}{2} \hat{S}_{ij} + \frac{mp_{(i} n S_{j)}}{np(m - np)}$$

$$\hat{p}_i = p_i + K_{ij} n S^j + A^{kl} e_{(j)k} e_{l,i}^{(j)} - \left(\frac{1}{2} S_{kj} + \frac{p_{(k} n S_{j)}}{np} \right) \Gamma^{kj}_i$$

- Matter Lagrangian density now, with $\hat{\Omega}^{(i)(j)} = \hat{\Lambda}_{[k]}^{(i)} \hat{\Lambda}^{[k](j)}$,

$$\mathcal{L}_M = A^{ij} e_{(k)i} e_{j,0}^{(k)} \hat{\delta} + \hat{p}_i \dot{\hat{z}}^i \hat{\delta} + \frac{1}{2} \hat{S}_{(i)(j)} \hat{\Omega}^{(i)(j)} \hat{\delta} - N \mathcal{H}^{\text{matter}} + N^i \mathcal{H}_i^{\text{matter}}$$



ADM Formalism with Spin

- Legendre transformation for gravitational field.
- Spatial symmetric gauge (Kibble 1963): $e_{(i)j} = e_{ij} = e_{ji}$

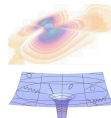
$$e_{ij}e_{jk} = \gamma_{ik} \quad \Rightarrow \quad (e_{ij}) = \sqrt{(\gamma_{ij})}$$

- Action:

$$W = \frac{1}{16\pi} \int d^4x \hat{\pi}^{ij} \gamma_{ij,0} + \int dt \left[\hat{p}_i \dot{z}^i + \frac{1}{2} \hat{S}_{(i)(j)} \hat{\Omega}^{(i)(j)} - H \right]$$
$$H = \int d^3x (N \mathcal{H} - N^i \mathcal{H}_i) + E[\gamma_{ij}]$$

- Canonical field momentum:

$$\hat{\pi}^{ij} = \pi^{ij} + 8\pi A^{(ij)} \hat{\delta} + 16\pi B_{kl}^{ij} A^{[kl]} \hat{\delta}$$
$$e_{k[i} e_{j]k,0} = B_{ij}^{kl} \gamma_{kl,0}$$



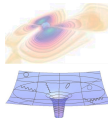
- Solve field constraints in ADMTT gauge.
- Fully reduced action:

$$W = \frac{1}{16\pi} \int d^4x \hat{\pi}_{\text{TT}}^{ij} h_{ij,0}^{\text{TT}} + \int dt \left[\hat{p}_i \dot{\hat{z}}^i + \frac{1}{2} \hat{S}_{(i)(j)} \hat{\Omega}^{(i)(j)} - H_{\text{ADM}} \right]$$
$$H_{\text{ADM}} = E[\hat{z}^i, \hat{p}_i, \hat{S}_{(i)(j)}, h_{ij}^{\text{TT}}, \hat{\pi}^{ij\text{TT}}]$$

- Fundamental equal-time Poisson brackets:

$$\{\hat{z}^i, \hat{p}_j\} = \delta_{ij}, \quad \{\hat{S}_{(i)}, \hat{S}_{(j)}\} = \varepsilon_{ijk} \hat{S}_{(k)}$$
$$\{h_{ij}^{\text{TT}}(\mathbf{x}, t), \hat{\pi}_{\text{TT}}^{kl}(\mathbf{x}', t)\} = 16\pi \delta_{ij}^{\text{TT}kl} \delta(\mathbf{x} - \mathbf{x}')$$

- Valid to all orders linear in spin.



Conserved Quantities

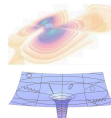
- Energy: $E = H_{\text{ADM}}$
- Total linear and angular momentum:

$$P_i = \sum_a \hat{p}_{ai} - \frac{1}{16\pi} \int d^3x \hat{\pi}_{\text{TT}}^{kl} h_{kl,i}^{\text{TT}}$$

$$J_{ij} = \sum_a (\hat{z}_a^i \hat{p}_{aj} - \hat{z}_a^j \hat{p}_{ai}) + \sum_a \hat{S}_{a(i)j)} \\ - \frac{1}{16\pi} \int d^3x (x^i \hat{\pi}_{\text{TT}}^{kl} h_{kl,j}^{\text{TT}} - x^j \hat{\pi}_{\text{TT}}^{kl} h_{kl,i}^{\text{TT}}) \\ - \frac{1}{16\pi} \int d^3x 2(\hat{\pi}_{\text{TT}}^{ik} h_{kj}^{\text{TT}} - \hat{\pi}_{\text{TT}}^{jk} h_{ki}^{\text{TT}})$$

- Boost: $J^{i0} \equiv K^i \equiv G^i - t P^i$
With center-of-mass vector:

$$G^i = -\frac{1}{16\pi} \int d^3\mathbf{x} x^i \Delta\phi$$



Stress-Energy-Tensor Algebra (Minkowski)

$$\{\mathcal{H}^m(x), \mathcal{H}^m(x')\} = -\mathcal{H}_i^m(x) \delta_{\mathbf{xx}',i} - \mathcal{H}_i^m(x') \delta_{\mathbf{xx},i}$$

$$\{\mathcal{H}_i^m(x), \mathcal{H}^m(x')\} = -\mathcal{H}^m(x) \delta_{\mathbf{xx}',i} - T_{ij}(x') \delta_{\mathbf{xx}',j}$$

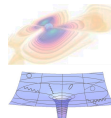
$$\{\mathcal{H}_i^m(x), \mathcal{H}_j^m(x')\} = -\mathcal{H}_j^m(x) \delta_{\mathbf{xx}',i} - \mathcal{H}_i^m(x') \delta_{\mathbf{xx}',j} + \partial_n \partial'_q [h_{inj}(x) \delta_{\mathbf{xx}'}]$$

$$h_{inj}(x) = \left[-\hat{S}_{q(n\mathcal{P}_i)j} - \delta^{kl} \frac{\hat{p}_k \hat{S}_{l(n\mathcal{P}_i)j\hat{p}_q}}{n\hat{p}(m-n\hat{p})} + \delta^{kl} \frac{\hat{p}_k \hat{S}_{l(q\mathcal{P}_j)(i\hat{p}_n)}}{n\hat{p}(m-n\hat{p})} \right] \hat{\delta}$$

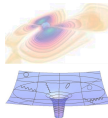
$$\mathcal{P}_{ij} \equiv \delta_{ij} - \frac{\hat{p}_i \hat{p}_j}{n\hat{p}^2}$$

$$\mathcal{H}^m(x) = T^{00}, \quad \mathcal{H}_i^m(x) = T^{0i}$$

- Local version of the Poincaré algebra.
- Minkowski limit of the gravitational constraint algebra.
- Dirac field also has $h_{inj}(x) \neq 0$.



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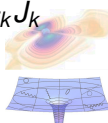
Order-by-Order Construction

- **Hamiltonian** $\hat{=}$ energy depending on canonical variables.
- Need to find canonical variables!
- Symmetries of the action:

$$P_i = \sum_a \hat{p}_{ai} + P_i^{\text{field}}, \quad J_i = \sum_a \left[\varepsilon_{ijk} \hat{z}_{aj} \hat{p}_{ak} + \hat{S}_{a(i)} \right] + J_i^{\text{field}}$$

- Global Poincaré algebra:

$$\begin{array}{lll} \{P_i, P_j\} = 0, & \{P_i, H\} = 0, & \{J_i, H\} = 0 \\ \{J_i, P_j\} = \varepsilon_{ijk} P_k, & \{J_i, J_j\} = \varepsilon_{ijk} J_k, & \{J_i, G_j\} = \varepsilon_{ijk} G_k \\ \{G_i, P_j\} = H \delta_{ij}, & \{G_i, H\} = P_i, & \{G_i, G_j\} = -\varepsilon_{ijk} J_k \end{array}$$



Canonical Variables at 2PN

- Calculate $\mathcal{H}_i^{\text{matter}}$:

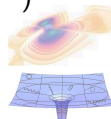
$$\mathcal{H}_i^{\text{matter}} = -\sqrt{\gamma} T_{i\nu} n^\nu$$

- Define canonical momentum \hat{p}_i as:

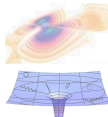
$$\hat{p}_i = \int d^3\mathbf{x} \mathcal{H}_i^{\text{matter}}$$

- Define spin $\hat{S}_{ij} = e_{i(k)} e_{j(l)} \varepsilon_{klm} \hat{S}_{(m)}$ such that $\hat{\mathbf{S}}^2 = \text{const.}$
- $\hat{\mathbf{z}}$ and $e_{i(k)}$ fixed by

$$J_{ij} = \hat{z}^i \hat{p}_j - \hat{z}^j \hat{p}_i + \varepsilon_{ijm} \hat{S}_{(m)} = \int d^3\mathbf{x} (x^i \mathcal{H}_j^{\text{matter}} - x^j \mathcal{H}_i^{\text{matter}})$$



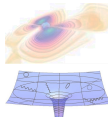
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Next-to-Leading Order (NLO) Spin-Orbit

First derived: Damour, Jaranowski, and Schäfer (2008)

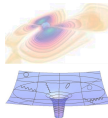
$$\begin{aligned} H_{\text{SO}}^{\text{NLO}} = & -\frac{((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^2} \left[\frac{5m_2 \hat{\mathbf{p}}_1^2}{8m_1^3} + \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2)}{4m_1^2} - \frac{3\hat{\mathbf{p}}_2^2}{4m_1 m_2} \right. \\ & \left. + \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})}{4m_1^2} + \frac{3(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})^2}{2m_1 m_2} \right] \\ & + \frac{((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^2} \left[\frac{(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2)}{m_1 m_2} + \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})}{m_1 m_2} \right] \\ & + \frac{((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{p}}_2)}{\hat{r}_{12}^2} \left[\frac{2(\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})}{m_1 m_2} - \frac{3(\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})}{4m_1^2} \right] \\ & - \frac{((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^3} \left[\frac{11m_2}{2} + \frac{5m_2^2}{m_1} \right] \\ & + \frac{((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12})}{\hat{r}_{12}^3} \left[6m_1 + \frac{15m_2}{2} \right] + (1 \leftrightarrow 2) \end{aligned}$$



NLO Spin₁-Spin₂

Partial result: Porto and Rothstein (2006)

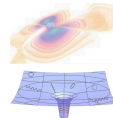
$$\begin{aligned}
 H_{S_1 S_2}^{\text{NLO}} = & \frac{1}{2m_1 m_2 \hat{r}_{12}^3} \left[\frac{3}{2} ((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) + \frac{1}{2} (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) \right. \\
 & + 6 ((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) - \frac{1}{2} (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_1) \\
 & - 15 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) + (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_2) \\
 & - 3 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) + 3 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) \\
 & + 3 (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) + 3 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) \\
 & \left. + 3 (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) - 3 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) \right] \\
 & + \frac{3}{2m_1^2 \hat{r}_{12}^3} \left[-((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_1 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) \right. \\
 & \left. + (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})^2 - (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) \right] \\
 & + \frac{3}{2m_2^2 \hat{r}_{12}^3} \left[-((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_2) \cdot \hat{\mathbf{n}}_{12}) ((\hat{\mathbf{p}}_2 \times \hat{\mathbf{S}}_1) \cdot \hat{\mathbf{n}}_{12}) \right. \\
 & \left. + (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12})^2 - (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) \right] \\
 & + \frac{6(m_1 + m_2)}{\hat{r}_{12}^4} \left[(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{S}}_2) - 2(\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_2 \cdot \hat{\mathbf{n}}_{12}) \right]
 \end{aligned}$$



$$\begin{aligned}
 \mathbf{G}_{\text{SO}}^{\text{NLO}} = & - \sum_a \frac{\hat{\mathbf{p}}_a^2}{8m_a^3} (\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \\
 & + \sum_a \sum_{b \neq a} \frac{m_b}{4m_a \hat{r}_{ab}} \left[((\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab}) \frac{5\hat{\mathbf{z}}_a + \hat{\mathbf{z}}_b}{\hat{r}_{ab}} - 5(\hat{\mathbf{p}}_a \times \hat{\mathbf{S}}_a) \right] \\
 & + \sum_a \sum_{b \neq a} \frac{1}{\hat{r}_{ab}} \left[\frac{3}{2} (\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_a) - \frac{1}{2} (\hat{\mathbf{n}}_{ab} \times \hat{\mathbf{S}}_a) (\hat{\mathbf{p}}_b \cdot \hat{\mathbf{n}}_{ab}) \right. \\
 & \quad \left. - ((\hat{\mathbf{p}}_b \times \hat{\mathbf{S}}_a) \cdot \hat{\mathbf{n}}_{ab}) \frac{\hat{\mathbf{z}}_a + \hat{\mathbf{z}}_b}{\hat{r}_{ab}} \right]
 \end{aligned}$$

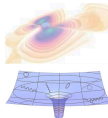
$$\mathbf{G}_{\text{S}_1 \text{S}_2}^{\text{NLO}} = \frac{1}{2} \sum_a \sum_{b \neq a} \left\{ \left[3(\hat{\mathbf{S}}_a \cdot \hat{\mathbf{n}}_{ab})(\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) - (\hat{\mathbf{S}}_a \cdot \hat{\mathbf{S}}_b) \right] \frac{\hat{\mathbf{z}}_a}{\hat{r}_{ab}^3} + (\hat{\mathbf{S}}_b \cdot \hat{\mathbf{n}}_{ab}) \frac{\hat{\mathbf{S}}_a}{\hat{r}_{ab}^2} \right\}$$

⇒ Poincaré algebra is fulfilled.



Outline

- 1 Introduction
- 2 Action Approach
- 3 Order-by-Order Construction
- 4 Results Linear in Spin
- 5 Higher Orders in Spin (Talk by S. Hergt, today 14:00)**



The Stress-Energy Tensor with Quadrupole

- Ansatz:

$$\sqrt{-g}T^{\mu\nu} = \int d\tau \left[t^{\mu\nu} \delta_{(4)} + (t^{\mu\nu\alpha} \delta_{(4)})_{||\alpha} + (t^{\mu\nu\alpha\beta} \delta_{(4)})_{||\alpha\beta} \right]$$

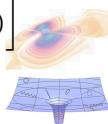
- Evaluate $T^{\mu\nu}_{||\nu} = 0$: (with D. Puetzfeld)

$$\frac{D(S^{\mu\nu})}{D\tau} = 2\rho^{[\mu} u^{\nu]} + \frac{4}{3} R^{\mu}_{\rho\beta\alpha} J^{\nu]\rho\beta\alpha}$$

$$\frac{D\rho_{\mu}}{D\tau} = -\frac{1}{2} R_{\mu\rho\beta\alpha} u^{\rho} S^{\beta\alpha} - \frac{1}{6} R_{\nu\rho\beta\alpha;\mu} J^{\nu\rho\beta\alpha}$$

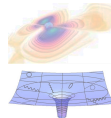
$$\sqrt{-g}T^{\mu\nu} = \int d\tau \left[u^{(\mu} \rho^{\nu)} \delta_{(4)} + \frac{1}{3} R^{\mu}_{\rho\beta\alpha} J^{\nu)\rho\beta\alpha} \delta_{(4)} + (u^{(\mu} S^{\nu)\alpha} \delta_{(4)})_{||\alpha} - \frac{2}{3} (J^{\mu\alpha\beta\nu} \delta_{(4)})_{||(\alpha\beta)} \right]$$

- Mass quadrupole $I^{\mu\nu}$: $J^{\nu\rho\beta\alpha} = -3u^{[\nu} I^{\rho][\beta} u^{\alpha]}$
- Spin-squared ansatz for $I^{\mu\nu}$. $\Rightarrow C_Q$ (= 1 for black holes)



NLO Spin₁-Spin₁ for Black Holes and Neutron Stars

$$\begin{aligned}
 H_{S_1^2}^{\text{NLO}} = & \frac{m_2}{m_1^3 \hat{r}_{12}^3} \left[\left(\frac{15}{4} - \frac{9}{2} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) + \left(\frac{5}{4} - \frac{5}{4} C_Q \right) \hat{\mathbf{p}}_1^2 \hat{\mathbf{S}}_1^2 \right. \\
 & + \left(-\frac{9}{8} + \frac{3}{2} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \hat{\mathbf{S}}_1^2 + \left(-\frac{21}{8} + \frac{9}{4} C_Q \right) \hat{\mathbf{p}}_1^2 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \\
 & \left. + \left(-\frac{5}{4} + \frac{3}{2} C_Q \right) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1)^2 \right] + \frac{C_Q}{m_1 m_2 \hat{r}_{12}^3} \left[\frac{9}{4} \hat{\mathbf{p}}_2^2 (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 - \frac{3}{4} \hat{\mathbf{p}}_2^2 \hat{\mathbf{S}}_1^2 \right] \\
 & + \frac{1}{m_1^2 \hat{r}_{12}^3} \left[\left(-\frac{3}{2} + \frac{9}{2} C_Q \right) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) \right. \\
 & + \left(-3 + \frac{3}{2} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) + \left(-\frac{3}{2} + \frac{9}{4} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) \hat{\mathbf{S}}_1^2 \\
 & + \left(\frac{3}{2} - \frac{3}{4} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) \hat{\mathbf{S}}_1^2 + \left(3 - \frac{21}{4} C_Q \right) (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{p}}_2) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \\
 & \left. - \frac{15}{4} C_Q (\hat{\mathbf{p}}_1 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{p}}_2 \cdot \hat{\mathbf{n}}_{12}) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 + \left(\frac{3}{2} - \frac{3}{2} C_Q \right) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_1) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{p}}_2) \right] \\
 & + \frac{m_2}{\hat{r}_{12}^4} \left[\left(2 + \frac{1}{2} C_Q \right) \hat{\mathbf{S}}_1^2 - \left(3 + \frac{3}{2} C_Q \right) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \right] \\
 & + \frac{m_2^2}{m_1 \hat{r}_{12}^4} \left[(1 + 2C_Q) \hat{\mathbf{S}}_1^2 - (1 + 6C_Q) (\hat{\mathbf{S}}_1 \cdot \hat{\mathbf{n}}_{12})^2 \right]
 \end{aligned}$$



Thank you for your attention

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