

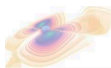
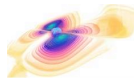
ADM canonical formulation with spin and application to post-Newtonian approximations

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DFG: SFB/TR7 Gravitational Wave Astronomy

ADM Hamiltonian and Global Poincaré Invariance

- Hamiltonian in ADMTT gauge (ADM Hamiltonian)
 $\hat{=}$ ADM energy depending on canonical variables:

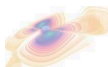
$$H_{\text{ADM}} = E[x_a^i, p_{ai}, h_{ij}^{\text{TT}}, \pi_{\text{TT}}^{jj}] = -\frac{1}{16\pi} \int d^3\mathbf{x} \Delta\phi$$
$$\gamma_{ij} = \left(1 + \frac{1}{8}\phi\right)^4 \delta_{ij} + h_{ij}^{\text{TT}}$$

- The algebra of global Poincaré invariance reads

$$\begin{aligned} \{P_i, P_j\} &= 0, & \{P_i, H\} &= 0, & \{J_i, H\} &= 0, \\ \{J_i, P_j\} &= \varepsilon_{ijk} P_k, & \{J_i, J_j\} &= \varepsilon_{ijk} J_k, & \{J_i, G_j\} &= \varepsilon_{ijk} G_k, \\ \{G_i, P_j\} &= H\delta_{ij}, & \{G_i, H\} &= P_i, & \{G_i, G_j\} &= -\varepsilon_{ijk} J_k, \end{aligned}$$

with

$$P_i = \sum_a p_{ai}, \quad J_i = \sum_a \left[\varepsilon_{ijk} \hat{z}_{a}^j p_{ak} + S_{a(i)} \right].$$

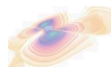


- Stress-energy tensor density in covariant SSC, $S^{\mu\nu} u_\nu = 0$:

$$\sqrt{-g} T^{\mu\nu} = \int d\tau \left[m u^\mu u^\nu \delta_{(4)} - (S^{\alpha(\mu} u^{\nu)}) \delta_{(4)} \right]_{||\alpha}$$
$$\delta_{(4)} \equiv \delta(x - z(\tau))$$

- EOM follow from $T^{\mu\nu}_{||\nu} = 0$:

$$\frac{DS^{\mu\nu}}{d\tau} = 0, \quad m \frac{Du_\mu}{d\tau} = \frac{1}{2} S^{\lambda\nu} u^\gamma R_{\mu\gamma\nu\lambda}^{(4)}$$



Identification of Canonical Variables

- Calculate $\mathcal{H}_i^{\text{matter}}$:

$$\mathcal{H}_i^{\text{matter}} = -\sqrt{\gamma} T_{iv} n^v$$

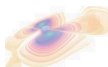
- Define canonical momentum p_i as:

$$p_i = \int d^3\mathbf{x} \mathcal{H}_i^{\text{matter}}$$

- Define spin $\hat{S}_{ij} = e_{i(k)} e_{j(l)} \varepsilon_{klm} S_{(m)}$ such that $\mathbf{S}^2 = \text{const.}$ and

$$J_{ij} = \hat{z}^i p_j - \hat{z}^j p_i + \varepsilon_{ijm} S_{(m)} = \int d^3\mathbf{x} (x^i \mathcal{H}_j^{\text{matter}} - x^j \mathcal{H}_i^{\text{matter}})$$

- Go over to canonical position variable \mathbf{z} by a Lie shift (such that one has the Newton-Wigner SSC in flat space).

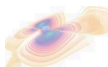


NLO Spin-Orbit Hamiltonian

First derived: Damour, Jaranowski, and Schäfer (2008)

$$\begin{aligned} H_{\text{SO}}^{\text{NLO}} = & -\frac{((\mathbf{p}_1 \times \mathbf{S}_1) \cdot \mathbf{n}_{12})}{r_{12}^2} \left[\frac{5m_2 \mathbf{p}_1^2}{8m_1^3} + \frac{3(\mathbf{p}_1 \cdot \mathbf{p}_2)}{4m_1^2} - \frac{3\mathbf{p}_2^2}{4m_1 m_2} \right. \\ & \left. + \frac{3(\mathbf{p}_1 \cdot \mathbf{n}_{12})(\mathbf{p}_2 \cdot \mathbf{n}_{12})}{4m_1^2} + \frac{3(\mathbf{p}_2 \cdot \mathbf{n}_{12})^2}{2m_1 m_2} \right] \\ & + \frac{((\mathbf{p}_2 \times \mathbf{S}_1) \cdot \mathbf{n}_{12})}{r_{12}^2} \left[\frac{(\mathbf{p}_1 \cdot \mathbf{p}_2)}{m_1 m_2} + \frac{3(\mathbf{p}_1 \cdot \mathbf{n}_{12})(\mathbf{p}_2 \cdot \mathbf{n}_{12})}{m_1 m_2} \right] \\ & + \frac{((\mathbf{p}_1 \times \mathbf{S}_1) \cdot \mathbf{p}_2)}{r_{12}^2} \left[\frac{2(\mathbf{p}_2 \cdot \mathbf{n}_{12})}{m_1 m_2} - \frac{3(\mathbf{p}_1 \cdot \mathbf{n}_{12})}{4m_1^2} \right] \\ & - \frac{((\mathbf{p}_1 \times \mathbf{S}_1) \cdot \mathbf{n}_{12})}{r_{12}^3} \left[\frac{11m_2}{2} + \frac{5m_2^2}{m_1} \right] \\ & + \frac{((\mathbf{p}_2 \times \mathbf{S}_1) \cdot \mathbf{n}_{12})}{r_{12}^3} \left[6m_1 + \frac{15m_2}{2} \right] + (1 \leftrightarrow 2) \end{aligned}$$

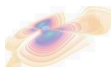
- Equivalent to Faye, Blanchet, Buonanno (2007)



NLO Spin₁-Spin₂ Hamiltonian

Partial result: Porto and Rothstein (2006)

$$\begin{aligned} H_{S_1 S_2}^{\text{NLO}} = & \frac{1}{2m_1 m_2 r_{12}^3} \left[\frac{3}{2} ((\mathbf{p}_1 \times \mathbf{S}_1) \cdot \mathbf{n}_{12}) ((\mathbf{p}_2 \times \mathbf{S}_2) \cdot \mathbf{n}_{12}) + \frac{1}{2} (\mathbf{S}_1 \cdot \mathbf{S}_2) (\mathbf{p}_1 \cdot \mathbf{p}_2) \right. \\ & + 6 ((\mathbf{p}_2 \times \mathbf{S}_1) \cdot \mathbf{n}_{12}) ((\mathbf{p}_1 \times \mathbf{S}_2) \cdot \mathbf{n}_{12}) - \frac{1}{2} (\mathbf{S}_1 \cdot \mathbf{p}_2) (\mathbf{S}_2 \cdot \mathbf{p}_1) \\ & - 15 (\mathbf{S}_1 \cdot \mathbf{n}_{12}) (\mathbf{S}_2 \cdot \mathbf{n}_{12}) (\mathbf{p}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_2 \cdot \mathbf{n}_{12}) + (\mathbf{S}_1 \cdot \mathbf{p}_1) (\mathbf{S}_2 \cdot \mathbf{p}_2) \\ & - 3 (\mathbf{S}_1 \cdot \mathbf{n}_{12}) (\mathbf{S}_2 \cdot \mathbf{n}_{12}) (\mathbf{p}_1 \cdot \mathbf{p}_2) + 3 (\mathbf{S}_1 \cdot \mathbf{p}_2) (\mathbf{S}_2 \cdot \mathbf{n}_{12}) (\mathbf{p}_1 \cdot \mathbf{n}_{12}) \\ & + 3 (\mathbf{S}_2 \cdot \mathbf{p}_1) (\mathbf{S}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_2 \cdot \mathbf{n}_{12}) + 3 (\mathbf{S}_1 \cdot \mathbf{p}_1) (\mathbf{S}_2 \cdot \mathbf{n}_{12}) (\mathbf{p}_2 \cdot \mathbf{n}_{12}) \\ & \left. + 3 (\mathbf{S}_2 \cdot \mathbf{p}_2) (\mathbf{S}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_1 \cdot \mathbf{n}_{12}) - 3 (\mathbf{S}_1 \cdot \mathbf{S}_2) (\mathbf{p}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_2 \cdot \mathbf{n}_{12}) \right] \\ & + \frac{3}{2m_1^2 r_{12}^3} \left[-((\mathbf{p}_1 \times \mathbf{S}_1) \cdot \mathbf{n}_{12}) ((\mathbf{p}_1 \times \mathbf{S}_2) \cdot \mathbf{n}_{12}) \right. \\ & \left. + (\mathbf{S}_1 \cdot \mathbf{S}_2) (\mathbf{p}_1 \cdot \mathbf{n}_{12})^2 - (\mathbf{S}_1 \cdot \mathbf{n}_{12}) (\mathbf{S}_2 \cdot \mathbf{p}_1) (\mathbf{p}_1 \cdot \mathbf{n}_{12}) \right] \\ & + \frac{3}{2m_2^2 r_{12}^3} \left[-((\mathbf{p}_2 \times \mathbf{S}_2) \cdot \mathbf{n}_{12}) ((\mathbf{p}_2 \times \mathbf{S}_1) \cdot \mathbf{n}_{12}) \right. \\ & \left. + (\mathbf{S}_1 \cdot \mathbf{S}_2) (\mathbf{p}_2 \cdot \mathbf{n}_{12})^2 - (\mathbf{S}_2 \cdot \mathbf{n}_{12}) (\mathbf{S}_1 \cdot \mathbf{p}_2) (\mathbf{p}_2 \cdot \mathbf{n}_{12}) \right] \\ & + \frac{6(m_1 + m_2)}{r_{12}^4} \left[(\mathbf{S}_1 \cdot \mathbf{S}_2) - 2(\mathbf{S}_1 \cdot \mathbf{n}_{12}) (\mathbf{S}_2 \cdot \mathbf{n}_{12}) \right] \end{aligned}$$



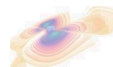
Hamiltonians from the Poincaré Algebra

Hergt and Schäfer (2008)

The full Hamiltonian up to 2PN enters the Poincaré algebra:

$$H = H_N + H_{1PN} + H_{2PN} + H_{SO}^{1PN} + H_{SO}^{2PN} + H_{S^2} + H_{S^3p} + H_{S^2p^2} + H_{S^4}$$

- Source terms in canonical variables sufficient for $H_{S_2^2 S_1 p_1}$, $H_{S_2^3 p_1}$, $H_{S_1^3 p_2}$, $H_{S_1^2 S_2 p_2}$, $H_{S_1^2 S_2^2}$, $H_{S_1 S_2^3}$, and $H_{S_2 S_1^3}$ were obtained from the Kerr-metric in ADM coordinates (HS 2007).
- Ansatzes for $H_{S_1^2 p^2}$, $H_{S_2^2 p^2}$, $H_{S_1^3 p_1}$, $H_{S_2^3 p_2}$, $H_{S_1^2 S_2 p_1}$, $H_{S_2^2 S_1 p_2}$, $H_{S_1^4}$, and $H_{S_2^4}$ are **fixed up to canonical transformation** by $\{G_i, H\} = P_i$.
- The static (linear momentum independent) part of the Hamilton constraint is needed to fix these remaining degrees of freedom.



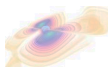
The Stress-Energy Tensor with Quadrupole

unpublished

- Stress-energy tensor density with quadrupole has the structure:

$$\sqrt{-g}T^{\mu\nu} = \int d\tau \left[t^{\mu\nu} \delta_{(4)} + (t^{\mu\nu\alpha} \delta_{(4)})_{||\alpha} + (t^{\mu\nu\alpha\beta} \delta_{(4)})_{||\alpha\beta} \right]$$

- Getting expressions for the $t^{\mu\nu\dots}$ from $T^{\mu\nu}_{||\nu} = 0$:
 - Dixon's work: Complicated definitions.
 - Tulczyjew's theorems: Complicated calculation.
- $t^{\mu\nu\alpha\beta}$ related to spin-squared mass quadrupole $I_1^{ij} \equiv \gamma^{ik} \gamma^{jl} \gamma^{mn} \hat{S}_{1km} \hat{S}_{1nl} + \frac{2}{3} \mathbf{S}_1^2 \gamma^{ij}$.



Ansatz for the Static Source Terms

$$\begin{aligned}
 \mathcal{H}_{S_1^2, \text{static}}^{\text{matter}} = & \frac{c_1}{m_1} \left(l_1^{ij} \delta_1 \right)_{;ij} + \frac{c_2}{m_1} R_{ij} l_1^{ij} \delta_1 + \frac{c_3}{m_1} \mathbf{S}_1^2 \left(\gamma^{ij} \delta_1 \right)_{;ij} + \frac{c_4}{m_1} \mathbf{R} \mathbf{S}_1^2 \delta_1 \\
 & + \frac{1}{8m_1} g_{mn} \gamma^{pj} \gamma^{ql} \gamma^{mi}_{,p} \gamma^{nk}_{,q} \hat{S}_{1ij} \hat{S}_{1kl} \delta_1 \\
 & + \frac{1}{4m_1} \left(\gamma^{jj} \gamma^{mn} \gamma^{kl}_{,m} \hat{S}_{1ln} \hat{S}_{1jk} \delta_1 \right)_{,i}
 \end{aligned}$$

- This ansatz is 3-dim. covariant, as p_i is not:

$$p_i = \int d^3 \mathbf{x} \mathcal{H}_i^{\text{matter}} = m v_i - \frac{1}{2} g_{ij} \gamma^{lm} \gamma^{kj}_{,m} \hat{S}_{kl} + \mathcal{O}(p^2) + \mathcal{O}(\hat{S}^2)$$

- Terms like $l_1^{ij}_{;k} \delta_1$ or $l_1^{ij} \delta_{1;k}$ can not appear.
- γ_{ij} for Kerr $\Rightarrow c_1 = -\frac{1}{2}$.
- Lapse function for Kerr $\Rightarrow c_2 = 0$.
- c_3 and c_4 do not contribute to the Hamiltonian.

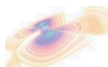


NLO Spin₁-Spin₁ Hamiltonian

$$\begin{aligned}
 H_{S_1^2}^{\text{NLO}} = & \frac{1}{r_{12}^3} \left[\frac{m_2}{4m_1^3} (\mathbf{p}_1 \cdot \mathbf{S}_1)^2 + \frac{3m_2}{8m_1^3} (\mathbf{p}_1 \cdot \mathbf{n}_{12})^2 \mathbf{S}_1^2 - \frac{3m_2}{8m_1^3} \mathbf{p}_1^2 (\mathbf{S}_1 \cdot \mathbf{n}_{12})^2 \right. \\
 & - \frac{3m_2}{4m_1^3} (\mathbf{p}_1 \cdot \mathbf{n}_{12}) (\mathbf{S}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_1 \cdot \mathbf{S}_1) - \frac{3}{4m_1 m_2} \mathbf{p}_2^2 \mathbf{S}_1^2 \\
 & + \frac{9}{4m_1 m_2} \mathbf{p}_2^2 (\mathbf{S}_1 \cdot \mathbf{n}_{12})^2 + \frac{3}{4m_1^2} (\mathbf{p}_1 \cdot \mathbf{p}_2) \mathbf{S}_1^2 \\
 & - \frac{9}{4m_1^2} (\mathbf{p}_1 \cdot \mathbf{p}_2) (\mathbf{S}_1 \cdot \mathbf{n}_{12})^2 + \frac{3}{4m_1^2} (\mathbf{p}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_2 \cdot \mathbf{n}_{12}) \mathbf{S}_1^2 \\
 & - \frac{3}{2m_1^2} (\mathbf{p}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_2 \cdot \mathbf{S}_1) (\mathbf{S}_1 \cdot \mathbf{n}_{12}) \\
 & + \frac{3}{m_1^2} (\mathbf{p}_2 \cdot \mathbf{n}_{12}) (\mathbf{p}_1 \cdot \mathbf{S}_1) (\mathbf{S}_1 \cdot \mathbf{n}_{12}) \\
 & \left. - \frac{15}{4m_1^2} (\mathbf{p}_1 \cdot \mathbf{n}_{12}) (\mathbf{p}_2 \cdot \mathbf{n}_{12}) (\mathbf{S}_1 \cdot \mathbf{n}_{12})^2 \right] \\
 & - \frac{m_2}{r_{12}^4} \left[\frac{9}{2} (\mathbf{S}_1 \cdot \mathbf{n}_{12})^2 - \frac{5}{2} \mathbf{S}_1^2 + \frac{7m_2}{m_1} (\mathbf{S}_1 \cdot \mathbf{n}_{12})^2 - \frac{3m_2}{m_1} \mathbf{S}_1^2 \right]
 \end{aligned}$$



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Higher Post-Newtonian Orders

with Han Wang, in preparation

- Need spin corrections to canonical field momentum:

$$\pi_{\text{can}}^{ij} = \pi_{\text{field}}^{ij} + \pi_{\text{spin}}^{ij},$$

$$\pi_{\text{field}}^{ij} = -\sqrt{\gamma}(\gamma^{ik}\gamma^{jl} - \gamma^{jj}\gamma^{kl})K_{kl}.$$

- Choose π_{spin}^{ij} such that:

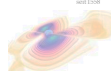
$$P_i = \sum_a p_{ai} - \frac{1}{16\pi} \int d^3x \pi_{\text{can}}^{k/TT} h_{kl,i}^{TT}$$

$$J_{ij} = \sum_a (\hat{z}_a^i p_{aj} - \hat{z}_a^j p_{ai}) + \sum_a \hat{S}_{a(i)(j)}$$

$$- \frac{1}{16\pi} \int d^3x (x^i \pi_{\text{can}}^{k/TT} h_{kl,j}^{TT} - x^j \pi_{\text{can}}^{k/TT} h_{kl,i}^{TT})$$

$$- 2 \frac{1}{16\pi} \int d^3x (\pi_{\text{can}}^{ikTT} h_{kj}^{TT} - \pi_{\text{can}}^{jkTT} h_{ki}^{TT})$$

- Got Hamiltonian for field evolution at formal 3.5PN.
- Checked 1PN energy flux (Kidder 1995).



- Action:

$$W[e_{a\mu}, z^\mu, \tilde{p}_\mu, \Lambda^{Ca}, S_{ab}, \lambda_1^a, \lambda_{2[i]}, \lambda_3] = \int d^4x \mathcal{L}$$
$$\Lambda^{Aa} \Lambda^{Bb} \eta_{AB} = \eta^{ab}$$

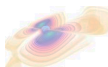
- Lagrangian densities $\mathcal{L} = \mathcal{L}_G + \mathcal{L}_M + \mathcal{L}_C$:

$$\mathcal{L}_M = \int d\tau \left[\left(\tilde{p}_\mu - \frac{1}{2} S_{ab} \omega_\mu^{ab} \right) \frac{dz^\mu}{d\tau} + \frac{1}{2} S_{ab} \frac{d\theta^{ab}}{d\tau} \right] \delta_{(4)}$$

$$\mathcal{L}_C = \int d\tau \left[\lambda_1^a \tilde{p}^b S_{ab} + \lambda_{2[i]} \Lambda^{[i]a} \tilde{p}_a - \frac{\lambda_3}{2} (\tilde{p}^2 + m^2) \right] \delta_{(4)}$$

$$\mathcal{L}_G = \frac{1}{16\pi} \sqrt{-g} R^{(4)}$$

$$d\theta^{ab} = \Lambda_C^a d\Lambda^{Cb}$$



Reduction of the Matter Part

in covariant SSC and time gauge

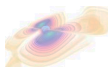
$$\mathcal{L}_M = \mathcal{L}_{MK} + \mathcal{L}_{MC} + \mathcal{L}_{GK}$$

- Kinetic matter part:

$$\begin{aligned} \mathcal{L}_{MK} = & \left[\tilde{p}_i + K_{ij} n S^j + A^{kl} e_{(j)k} e_{l,i}^{(j)} - \left(\frac{1}{2} S_{kj} + \frac{\tilde{p}_{(k} n S_{j)}}{n \tilde{p}} \right) \Gamma^{kj} \right] \dot{z}^i \delta \\ & + \frac{n S^i}{2 n \tilde{p}} \dot{\tilde{p}}_i \delta + \left[S_{(i)(j)} + \frac{n S_{(i)} \tilde{p}_{(j)} - n S_{(j)} \tilde{p}_{(i)}}{n \tilde{p}} \right] \frac{\Lambda_{[k}^{(i)} \dot{\Lambda}^{(j)k]}}{2} \delta \end{aligned}$$

- Constraint part $\mathcal{L}_{MC} = N \mathcal{H}^{\text{matter}} - N^i \mathcal{H}_i^{\text{matter}}$:

$$\mathcal{H}^{\text{matter}} = \sqrt{\gamma} T_{\mu\nu} n^\mu n^\nu, \quad \mathcal{H}_i^{\text{matter}} = -\sqrt{\gamma} T_{i\nu} n^\nu$$



Reduction of the Field Part and Result

- Transition to (generalized) Newton-Wigner variables.

$$\Rightarrow \hat{\mathcal{L}}_{GK} = \hat{A}^{ij} e_{(k)i} e_{(k)j,0} \delta, \quad g_{ik} g_{jl} \hat{A}^{kl} = \frac{1}{2} \hat{S}_{ij} + \frac{mp_{(i} n S_{j)}}{np(m - np)}$$

- Spatial symmetric gauge (Kibble 1963): $e_{(i)j} = e_{ij} = e_{ji}$

$$e_{ij} e_{jk} = g_{ik} \quad \Rightarrow \quad e_{ij} = \sqrt{g_{ij}}$$

- Definition of field momentum:

$$\pi_{\text{can}}^{ij} = \pi^{ij} + 8\pi \hat{A}^{(ij)} \delta + 16\pi B_{kl}^{ij} \hat{A}^{[kl]} \delta$$

$$e_{k[i} e_{j]k,0} = B_{ij}^{kl} g_{kl,0}$$

- Result:

$$W = \frac{1}{16\pi} \int d^4 x \pi_{\text{can}}^{ij\text{TT}} h_{ij,0}^{\text{TT}} + \int dt \left[p_i \dot{z}^i + \frac{1}{2} \hat{S}_{(i)(j)} \frac{d\hat{\theta}^{(i)(j)}}{dt} - E \right]$$



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Thank you for your attention!

